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# Ground-based remote sensing of nitrous oxide (N<sub>2</sub>O) over Hefei, eastern China from high-resolution solar spectra

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## ABSTRACT

We for the first time demonstrate ground-based remote sensing of Nitrous Oxide (N<sub>2</sub>O) over Hefei in eastern China from high resolution Fourier Transform Infra-Red (FTIR) solar spectra. We have retrieved Column-averaged Abundance of N<sub>2</sub>O ( $X_{N_2O}$ ) from both Near-Infrared (NIR, 4000 to 11,000 cm<sup>-1</sup>) and Mid-Infrared (MIR, 2400 to 3200 cm<sup>-1</sup>) solar spectra and inspected their agreement. Generally, NIR and MIR measurements agree well with a correlation coefficient of 0.86 and an average difference of (1.33 ± 4.05) ppbv (NIR – MIR). By correcting the bias of the two datasets, we combine the NIR and MIR measurements to investigate seasonality and inter-annual trend of  $X_{N_2O}$  over Hefei. The observed monthly mean time series of  $X_{N_2O}$  minimize in June and maximize in September, with values of (316.55 ± 12.22) ppbv and (322.05 ± 12.93) ppbv, respectively. The  $X_{N_2O}$  time series from 2015 to 2020 showed an inter-annual trend of (0.53 ± 0.10) %/year over Hefei, China. We also compared the FTIR  $X_{N_2O}$  observations with GEOS-Chem model  $X_{N_2O}$  simulations. They are in reasonable agreement with a correlation coefficient (R) of 0.71, but GEOS-Chem model underestimated the seasonality of the observations. This study can enhance current knowledge of ground-based high-resolution FTIR remote sensing of N<sub>2</sub>O in the atmosphere and contribute to generating a new reliable N<sub>2</sub>O dataset for climate change research.

## ARTICLE HISTORY

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## KEYWORDS

Greenhouse Gases (GHGs); remote sensing; high resolution; Fourier Transform Infra-Red (FTIR)

## 1. Introduction

N<sub>2</sub>O ranks as the third largest contributor to global warming (Stein and Yung 2003). In stratospheric atmosphere, N<sub>2</sub>O is a significant anthropogenic driver of ozone depletion and a major source of Nitric Oxide (NO) (Nevison et al. 2011; Portmann, Daniel, and Ravishankara 2012; Ravishankara, Daniel, and Portmann 2009). N<sub>2</sub>O is a very stable Greenhouse Gas (GHG) and it has a lifetime of about 114 years in the atmosphere (United States Environmental Protection Agency 2019). Globally averaged N<sub>2</sub>O in 2016 relative to 1750 has increased by 22%, corresponding to an inter-annual increase trend of 0.7 to 1.0 ppbv/year (Machida et al. 1995; Prinn et al. 1990, 2000). On global scale, the source of N<sub>2</sub>O is attributed to natural emissions by 60% and anthropogenic emissions by 40%. The sink of N<sub>2</sub>O is mainly depleted by ultraviolet photolysis in the stratosphere (Morgan et al. 2004). Additional sink of N<sub>2</sub>O is through photo oxidation with O(<sup>1</sup>D) to produce Nitrogen Oxides (NO<sub>x</sub>) (Montzka, Dlugokencky, and Butler 2011). The Kyoto Protocol has included N<sub>2</sub>O as one of the major GHGs targeted for emissions reduction (UNFCC 1997). Precise detection of the abundance

and variability of N<sub>2</sub>O is of great significance not only for understanding its source, transport, chemical mechanism, and thus establishing a scientific link between its variability and climate change, but also for improving theoretical Chemical Transport Models (CTMs) and potential control purposes.

Compared with in situ surface data, column observations pass through a larger portion of the atmosphere, resulting in a less temporal variability but still keeping information on surface fluxes (Wunch et al. 2011; Yin et al. 2019). As a result, the horizontal gradients of column observations are more associated with sources and sinks of underlying regional fluxes (Sun et al. 2020, 2018a, 2021a, 2021b). Ground-based high-resolution FTIR spectroscopy has been verified as one of the most precise remote sensing tools to derive total columns and Vertical Mixing Ratio (VMR) profile of N<sub>2</sub>O. Both the Total Columns Carbon Observing Network (TCCON) and Network for Detection of Atmospheric Composition Change (NDACC) networks use high resolution FTIR spectrometers to monitor atmospheric N<sub>2</sub>O abundance. The TCCON/NDACC networks initiated in 2004/1992 and a few stations have provided datasets of total columns and VMR profile of N<sub>2</sub>O for more than 20

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years. The California Laboratory for Atmospheric Remote Sensing (CLARS) which is not afflicted with the TCCON/NDACC networks also uses high resolution FTIR spectrometer to observe the long-term variability of  $N_2O$  over the Los Angeles basin (Addington et al. 2021). The inter-annual trends of  $N_2O$  have been reported in different regions with ground-based high resolution FTIR observations. Angelbratt et al. (2011) estimated that  $N_2O$  inter-annual trends varied between  $(0.19 \pm 0.01)$  and  $(0.40 \pm 0.02)$  %/year in four European FTIR stations during 1996–2007. Zhou et al. (2019) have compared the abundance and inter annual trends of  $N_2O$  from the TCCON and NDACC observations at seven FTIR observation sites around the globe, and concluded that  $N_2O$  in these sites varied over  $(0.81 \pm 0.04)$  to  $(0.93 \pm 0.02)$  %/year during 2007–2017. Sagar et al. (2022) also retrieved  $CO_2$ ,  $CH_4$  and  $CO$  concentrations using ground-based EM27/SUN FTIR in India.

Historically, most FTIR sites are distributed in Northern America and Europe, but the number of sites in other parts of the world is very sparse. Currently, only two TCCON stations have been set up in China, the Hefei station in eastern China ( $117.2^\circ E$ ,  $32.0^\circ N$ ) and the Xianghe station in northern China ( $116.96^\circ E$ ,  $39.75^\circ N$ ) (Tian et al. 2017; Wang et al. 2017; Yang et al. 2020; Yin et al. 2021a). In this manuscript, we first retrieve and characterize the variability of atmospheric  $N_2O$  over the Hefei station by using high-resolution FTIR spectroscopy. The  $N_2O$  retrievals derived from high resolution NIR (4,000 to 11,000  $cm^{-1}$ ) and MIR (2,400 to 3,200  $cm^{-1}$ ) solar spectra are compared. Furthermore, the NIR and MIR measurements from 2015 to 2020 are combined to investigate seasonality and inter annual trend of  $X_{N_2O}$  over Hefei. In addition, a comparison between GEOS-Chem simulations and FTIR observations are also performed. This study can enhance current knowledge of ground-based high-resolution FTIR remote sensing of  $N_2O$  in the atmosphere and contribute to generate a new reliable  $N_2O$  dataset for climate change research.

## 2. Methodology

### 2.1. FTIR observation

#### 2.1.1. Instrumentation

The operational FTIR observation site ( $117.2^\circ E$ ,  $31.5^\circ N$ , 30 m a.s.l. (above sea level)), run by the Anhui Institute of Optics and Fine Mechanics, Chinese Academy of Sciences (AIOFM-CAS), is located in the suburb of Hefei city (the capital city of Anhui Province) in central-eastern China. The downtown Hefei is located to the southeast of this site and is densely populated with eight million people. The site is surrounded by wetlands or cultivated lands in other

directions. The anthropogenic emissions mainly come from the city and the natural emissions are originated from cultivated lands or wetlands. Routine observations at the site have been implemented since 2014. The observatory is currently the only site in China that has a continuously-operating solar FTIR, making it crucial to calibrate and validate the satellite data or model simulations in this important region.

The observatory at Hefei includes a high-resolution FTIR spectrometer, a solar tracker, and a weather station. The FTIR spectrometer is installed inside a room, while the solar tracker and the weather station are placed on the roof of the room. The IFS125HR spectrometers show outstanding capabilities such as the accuracy and stability, and have been extensively applied within the TCCON/NDACC networks (Sun et al. 2021a, 2021b; Yin et al. 2020, 2021a, 2022b). The IFS125HR spectrometer at the Hefei site includes four beam splitters, nine optical path compartments, and six detectors. This spectrometer covers a wide wavenumber range of 400 to 50,000  $cm^{-1}$  with a maximum spectral resolution of 0.001  $cm^{-1}$ . With the state-of-the-art Camtracker mode, the solar tracker captures and guides solar beam uninterruptedly into the FTIR spectrometer with a precision of  $\pm 0.1$  mrad (Hall et al. 2011). The meteorological station includes sensors for relative humidity ( $\pm 3.0\%$ ), air temperature ( $\pm 0.30^\circ C$ ), air pressure ( $\pm 0.1$  hpa), solar radiation ( $\pm 5.0\%$ ), wind speed ( $\pm 0.40$  m/s), and wind direction ( $\pm 4.0^\circ$ ) (Sun et al. 2018a, 2017). These meteorological data are applied for retrievals or source attribution of the measurements.

In present work, we equip the FTIR spectrometer with a  $CaF_2$  beam splitter and an InGaAs detector for  $N_2O$  observations in the NIR range (4000–11,000  $cm^{-1}$ ) and a KBr beam splitter, a filter with a center wavenumber of 2800  $cm^{-1}$ , and an InSb detector for  $N_2O$  observations in the MIR range (2400–3100  $cm^{-1}$ ). We restricted the entrance aperture size to be 1 mm for all NIR measurements, while a size varying from 0.10 to 2.5 mm was implemented for the MIR measurements. Entrance aperture implemented in this manner can maximize the signal-to-noise ratio (SNR) while retaining unsaturated signals. The NIR and MIR solar spectra are saved alternatively during routine operations. The number of NIR and MIR  $N_2O$  observations within a sunny day range from 1 to 48.

#### 2.1.2. Retrieval and characterization of $N_2O$ in the NIR range

We utilize the GGG version 2020 to retrieve  $N_2O$  total columns from NIR solar spectra. GGG (Wunch et al. 2011) is a TCCON standard software which operates a suite of software package to retrieve total columns of various gaseous constituents from FTIR solar spectra. In order to obtain a site-to-site consistency of the retrievals, a same version of software is mandatorily

used for all TCCON sites (Wunch et al. 2015). The GGG first operates a subroutine procedure of interferogram-to-spectrum (I2S) to convert the interferograms into spectra. A subroutine procedure of “GSETUP” is then operated to generate *a priori* profiles of N<sub>2</sub>O and other trace gases. Meanwhile, *a priori* profiles of temperature, pressure, and H<sub>2</sub>O are straightly interpolated from the NCEP/NCAR reanalysis. This subroutine procedure takes into account a series of issues, including an interhemispheric gradient, a secular increase, stratospheric decay, and seasonal cycle. Finally, a subroutine procedure called the nonlinear least-squares fitting algorithm “GFIT” scales iteratively the *a priori* profiles to produce calculated spectra until the best fit to the measured spectra. The total columns of N<sub>2</sub>O ( $T_{N_2O}$ ) is thus calculated via Equation (1).

$$T_{N_2O} = \int_{z_s}^{z_{top}} A_m(z) \times f_{N_2O}(z) dz, \quad (1)$$

where  $f_{N_2O}(z)$  represents the integral of the mole fraction of N<sub>2</sub>O,  $A_m(z)$  is air-mass profile. The upper and lower limits of the integration,  $z_{top}$  and  $z_s$ , denote the top of atmosphere (TOA) and the surface, respectively. Unlike total columns which is affected by topography and surface pressure, column-averaged abundance has a low sensitivity to fluctuations of surface pressure and H<sub>2</sub>O. This characteristic allows direct comparisons of column observations with in situ data in different seasons and thus is advantageous for atmospheric cycle investigations (Wunch et al. 2015).  $T_{N_2O}$  can be converted into column-averaged abundance of N<sub>2</sub>O ( $X_{N_2O}$ ) via Equation (2). Taking the ratio in Equation (2) minimizes the systematic errors which are common to N<sub>2</sub>O and O<sub>2</sub>.

$$X_{N_2O} = \frac{T_{N_2O}}{T_{air}} = \frac{T_{N_2O}}{T_{O_2}} \times 0.2095 \quad (2)$$

where  $T_{air}$  and  $T_{O_2}$  are total columns of dry air and O<sub>2</sub>, respectively. Input parameters and setups for N<sub>2</sub>O and O<sub>2</sub> retrievals from NIR spectra using GGG version 2020 are summarized in Table 1 (Kalnay et al. 1996). As described in Wunch et al. (2011) and Wunch et al. (2015), N<sub>2</sub>O is retrieved in three spectral micro windows (MWs: 4373.5–4416.9 cm<sup>-1</sup>, 4418.55–4441.65

cm<sup>-1</sup> and 4682.95–4756.05 cm<sup>-1</sup>), and O<sub>2</sub> is retrieved in the spectral window of 7765.0–8004.0 cm<sup>-1</sup>. The cross-interferences of H<sub>2</sub>O, CO<sub>2</sub>, and HF are considered in the O<sub>2</sub> window and the cross-interferences of CH<sub>4</sub>, H<sub>2</sub>O, HDO and CO<sub>2</sub> are considered in the N<sub>2</sub>O window. Spectroscopic parameters of all gases are extracted from the HITRAN 2020 database (Hill et al. 2013; Rothman et al. 2009).

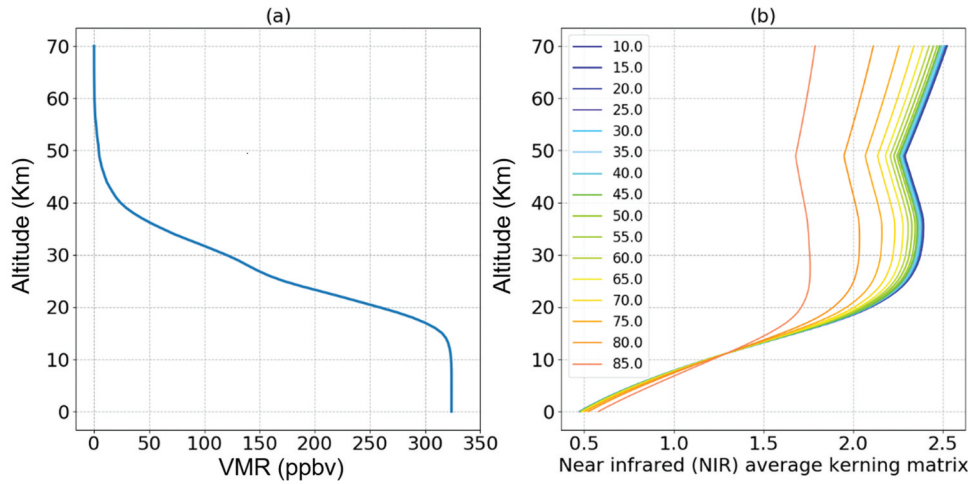
We collect and retrieve NIR spectra at Hefei following the TCCON recommendation, and thus the error budget for N<sub>2</sub>O by the TCCON community could also apply to Hefei observations. Wunch et al. (2011) perturbed each potential error by a realistic amount in the GGG forward model and calculated the relative difference in  $X_{N_2O}$  relative to the unperturbed case. This sensitivity study was proceeded with solar spectra collected on a clear day at the Lamont site, which spanned a large difference in H<sub>2</sub>O, temperature, and solar zenith angle (SZA). The total error is estimated as the sum in quadrature of all potential errors. According to Wunch et al. (2015), the total errors of  $X_{N_2O}$  are ~1% and reasonably independent of SZA. The dominant error sources are shear misalignment, *a priori* profile shape, observer-sun Doppler stretch (OSDS), and zero level offsets. At high SZAs, the *a priori* temperature profile and pointing offsets become significant. The *a priori* profile and the column AVK for  $X_{N_2O}$  retrieval at Hefei in the NIR range are presented in Figure 1(a,b), respectively. The results show that the sensitivity of  $X_{N_2O}$  retrieval at Hefei in the NIR range decreases as the SZAs increases, but all sensitivities are greater than 0.5 between surface and 70 km. It means that more than 50% of retrieval information comes from observation, i.e. the *a priori* information affects the retrieval by less than 50% in both troposphere and stratosphere. It also indicates that the NIR observations for N<sub>2</sub>O at Hefei are sensitive to both troposphere and stratosphere regardless of SZAs.

### 2.1.3. Retrieval and characterization of N<sub>2</sub>O in the MIR range

We use the SFIT4 software to retrieve N<sub>2</sub>O from MIR spectra. Inputs and setups of the SFIT4 follow the

**Table 1.** Input parameters and setups for N<sub>2</sub>O and O<sub>2</sub> retrievals from NIR solar spectra.

Species	N <sub>2</sub> O	O <sub>2</sub>
Retrieval Code	GGG2020	GGG2020
Spectroscopy	HITRAN 2020	HITRAN 2020
A priori profiles of temperature and pressure	NCEP	NCEP
A priori profiles of gases	GGG2020 code	GGG2020 code
Micro windows (cm <sup>-1</sup> )	4373.5–4416.9 4418.55–4441.65 4682.95–4756.05	7765–8005
Micro spectral window width (cm <sup>-1</sup> )	43.4 23.1 73.1	240
Retrieved interfering species	CH <sub>4</sub> , H <sub>2</sub> O, CO <sub>2</sub> , HDO	CO <sub>2</sub> , H <sub>2</sub> O, HF



**Figure 1.** (a) The a priori profile of  $N_2O$  used in NIR spectra retrieval. (b) Averaging kernel (AVK) matrix for  $N_2O$  retrieval with GGG2020.

NDACC recommendation and are tabulated in Table 2. We take monthly mean profiles of WACCM model simulations between 1980 and 2020 as the *a priori* profiles of all gases (not including  $H_2O$ ). The *a priori* profiles of meteorological parameters such as Temperature, humidity, and Pressure are straightly taken from NCEP/NCAR reanalysis (Kalnay et al. 1996). Three MWs of  $2441.8\text{--}2444.6\text{ cm}^{-1}$ ,  $2481.2\text{--}2482.5\text{ cm}^{-1}$  and  $2806.05\text{--}2806.55\text{ cm}^{-1}$  were used to retrieve  $N_2O$  profiles.  $CH_4$ , HDO, and  $H_2O$  show absorption interference with  $N_2O$  in all retrieval windows. In addition to  $N_2O$  profile, we also retrieve total columns of  $CH_4$  and HDO, and  $H_2O$  profile in the same MWs to minimize the cross absorption interferences. The spectroscopic parameters of all gases are extracted from the HITRAN 2020 database, which is same as the NIR spectra retrievals (Hill et al. 2013; Rothman et al. 2009).

To regularizing the retrieval, we set the diagonal values of the noise covariance matrix  $S_\epsilon$  to be square inverse of the SNR of each spectrum, and its non-diagonal values to be zero. We set the diagonal values of *a priori* covariance matrices  $S_a$  to be covariance of WACCM simulations between 1980 and 2020, and their non-diagonal values to be zero. In order to avoid inconsistencies in the total columns due to optical misalignment, we use the true instrumental line shape (ILS) derived from regular low-pressure HBr cell measurements to retrieve the time series of  $N_2O$ . (Hase 2012; Hase, Blumenstock, and Paton-Walsh 1999; Sun et al. 2018a, 2018b).

The AVK matrix  $A$  characterizes vertical information of the retrieval. The trace of  $A$  is called the degrees of freedom for signal (DOFS), which indicates that how much independent information can be derived

**Table 2.** Inputs and setups for  $N_2O$  retrieval from MIR spectra at Hefei.

Gases	$N_2O$	
Code	SFIT4	
A priori profiles for all gases (not including $H_2O$ )	Statistics of WACCM simulations	
A priori profiles of Temp., Humid., and Pres.	NCEP	
Spectroscopy	HITRAN2016	
Micro-windows ( $\text{cm}^{-1}$ )	2441.8–2444.6 2481.2–2482.5 2806.05–2806.55	
Micro spectral window width ( $\text{cm}^{-1}$ )	2.8 1.3 0.5	
Retrieved interfering gases	$CH_4$ , $CO_2$ , HDO	
Regularization	$S_\epsilon$	Real SNR (calculated in the actual observation in each spectrum)
	$S_a$	Covariance of WACCM simulation
ILS	measured	
Error (2.20%)	Random error: (1.18%) -measurement error (measurement) -interference errors: interfering gases (interfering_species), retrieval parameters (retrieval_parameters) -other errors: zero level (zshift), Temp. (temperature) Systematic error: (1.86%) -smoothing error (smoothing) -other errors: optical path difference (max_opd), spectral curvature (curvature), solar strength (solstrnth), field of view of incident solar beam (omega), shift of solar line (solshft), background slope (slope), solar zenith angle (sza), phase (phase), pressure induced line broadening (linepair_gas), temperature induced line broadening (linetair_gas), intensity of spectroscopic line (lineint_gas)	



from the retrieval. The area of **A** represents the retrieval sensitivity at each layer. **Figure 2** demonstrates the a priori and retrieved VMR profiles of  $N_2O$  (a), cumulative sum of DOFS (b), and the AVKs (c) for  $N_2O$  retrieval randomly selected in the MIR range. Ground-based FTIR  $N_2O$  retrievals in the MIR range at Hefei has a DOFS of 2.0 from surface to 12 km and 2.1 from 12 km to 40 km, which means that the MIR retrievals are sensitive in both troposphere and stratosphere. Typical DOFS over the total atmosphere for  $N_2O$  retrievals in the MIR range is 4.1, indicating that we can obtain almost four independent information on the retrieved VMR profile. Therefore, the  $N_2O$  total columns and  $X_{N_2O}$  discussed in this study are reliable. The  $N_2O$  total columns is obtained by integrating the  $N_2O$  profile from ground to the TOA. The  $X_{N_2O}$  is then obtained by a weighting average of the air mass and the  $N_2O$  profile from ground to the TOA. **Figure 2** shows that the retrieved profile deviates slightly from the a priori profile and is weighted for both troposphere and stratosphere due to the higher DOFS.

Error analysis for  $N_2O$  retrievals in the MIR range at Hefei follows the methodology of Rodgers (2000). We classified each individual error source as random or systematic error depending on if it varies randomly or is constant over observation. In present work, the random errors include  $z$ -shift error, measurement noise, and uncertainties of interfering species temperature, and retrieval parameters. The systematic errors include uncertainties of optical path difference, background curvature, solar line shift, the field of view of the incident solar beam, interferogram phase, the strength of solar line, pressure induced line broadening, temperature induced line broadening, and spectroscopic line intensity. **Table 2** concludes the random, systematic,

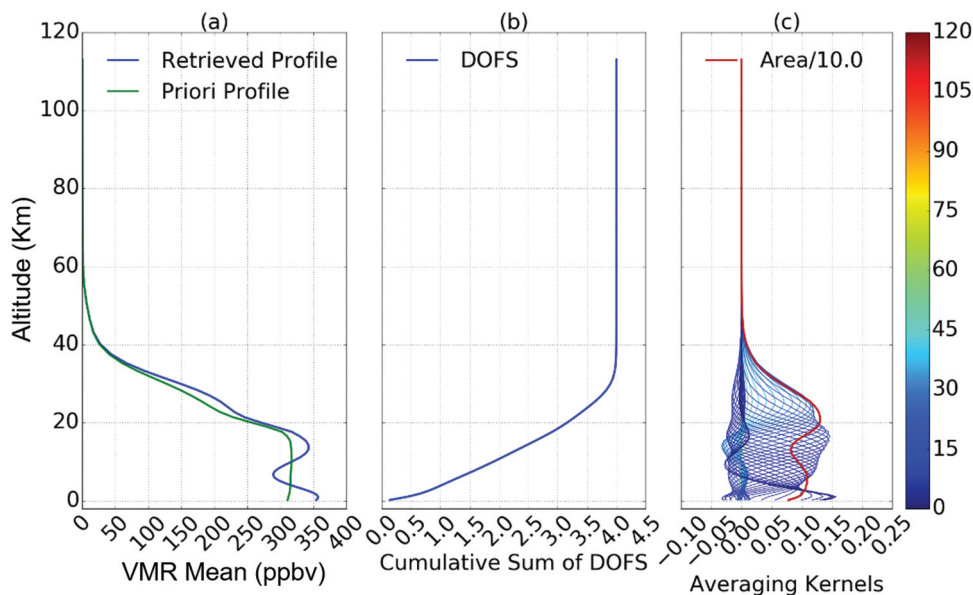
and total error for the  $N_2O$  retrieval exemplified in **Figure 3**. The input covariance matrix of observation noise was set to be the inverse square of the SNR of each spectrum and the temperature was estimated to be 2–7 K in the profile. The input covariance matrix of each interfering species was based on the WACCM climatology between 1980 and 2020. The input covariance matrix of field of view, background curvature, interferogram phase, optical path difference, shift of solar line, and strength of solar line were estimated to be 0.1%. The input covariance matrix of smoothing error and each retrieval parameter were obtained straightly from SFIT4 output. Spectroscopic uncertainties of spectroscopic line intensity, temperature, and pressure induced broadening coefficients for  $N_2O$  in the HITRAN 2020 line-list database were estimated to be 2%.

As shown in **Figure 3**, for  $N_2O$  retrieval in the MIR range at Hefei, the largest random error and systematic error are smoothing uncertainty (1.12%) and line intensity uncertainty (1.66%), respectively. Total error estimated as the sum in quadrature of total random errors (1.18%) and systematic errors (1.86%) is 2.20%.

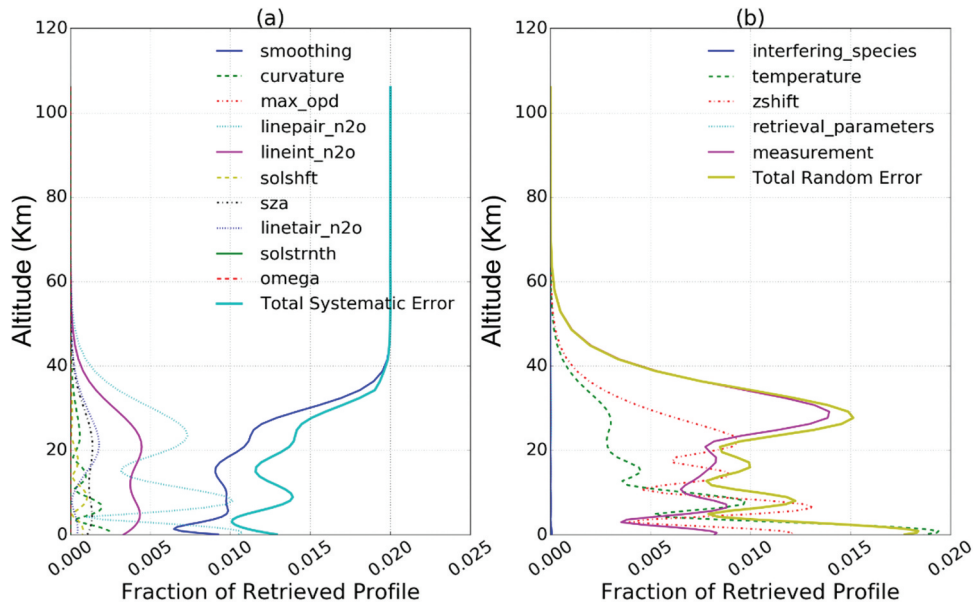
#### 2.1.4. Data filter criteria

We established a series of filter criteria to exclude the outliers that are heavily impacted by unstable weather conditions. Observations satisfying all the following criteria were regarded as valid observations and were applied in subsequent analysis, which excluded 23.4% of total observations.

- (1) Auxiliary data such as solar flux, surface temperature, humidity, and pressure must be collected synchronously with the observations.



**Figure 2.** (a) The a priori and retrieved profiles of  $N_2O$ , (b) cumulative sum of DOFS for  $N_2O$  retrieval, (c) the AVK matrix **A** and their area. The demonstrations are randomly selected  $N_2O$  retrieval in the MIR range.



**Figure 3.** (a) Systematic error budgets for  $N_2O$  retrievals in the MIR range. (b) Same as (a) but for random error budgets. Descriptions of the error components inside each panel are listed in Table 2.

- (2) The signal intensities of NIR spectra much vary between 7,000 and 22,000 counts and those of MIR spectra must vary between 10,000 and 20,000 counts. These thresholds well balanced the SNR and the non-linearity in the detector.
- (3) The observation scenario must be nearly cloud free. We reject any spectra which were saved when the variation of solar intensity is larger than 10%.
- (4) For both NIR and MIR solar spectra, the SZA must be smaller than  $80^\circ$ .
- (5) All retrievals must be converged and the concentrations at all vertical heights must be positive. For MIR spectra retrievals, the total DOFs must be greater than 3.0 to make sure that the retrieval at each partial layer comes more from observation.
- (6) The root means square error (RMS) of the fitting residual has to be smaller than 3% for  $N_2O$  retrieval in the NIR range and 5% in the NIR range.

## 2.2. GEOS-Chem model

We use GEOS-Chem model v12.2.1 to simulate  $N_2O$  around the globe at a  $2^\circ$  latitude  $\times$   $2.5^\circ$  longitude spatial resolution and 72 vertical pressure levels. We refer to Yin et al. (2019) for detailed description of model configuration. Briefly, the GEOS-Chem model is operated in a standard mode and driven by the GEOS-FP meteorological field. In order to drive the GEOS-Chem model, the GEOS-FP with an original spatial resolution of  $0.25^\circ \times 0.3125^\circ$  has been downgraded to  $2^\circ \times 2.5^\circ$ . The meteorological data has a vertical pressure resolution of 72 levels starting from ground to 0.015 hPa, and

temporal resolutions of 1 h for surface meteorological elements and 3 h for other meteorological elements.

Global anthropogenic emissions of aerosol species (mainly black carbon and organic carbon) and gases ( $SO_2$ ,  $NH_3$ ,  $NO_x$ , CO,  $CH_4$ , NMVOCs, etc.) between 2015 and 2017 are based on the MIX inventory over Asia and the CEDS (Community Emissions Data System) inventory in other parts of the world (Hoesly et al. 2018). Biomass burning emissions of atmospheric species are coupled with the GFED4 (Global Fire Emission Database version 4) inventory (Giglio, Randerson, and Van Der Werf 2013). The UCX (Universal tropospheric-stratospheric Chemistry eXtension) chemical mechanism and the FAST-JX v7.0 photolysis algorithm are implemented, which includes a detailed “ $NO_x$ - $O_x$ -hydrocarbon-aerosols” chemistry in the troposphere (Eastham, Weisenstein, and Barrett 2014). In stratosphere, bromine gases concentrations are based on climatology (Mclinden et al. 2000), ozone concentrations are estimated by the linearized ozone (Linoz) parameterization (Parrish et al. 2014), and other species concentrations are calculated by using archived monthly mean production s and loss rate from NASA’s GMI scheme (Murray et al. 2012; Rothman et al. 2009).

## 3. Results and discussion

### 3.1. Comparisons between NIR and MIR observations

We compare  $N_2O$  retrievals in the NIR and MIR ranges at Hefei to inspect the agreement of these two datasets and to determine how they can be combined to evaluate the variability of  $N_2O$ . The

NIR and MIR solar spectra in present work were saved with the same instrument, but the optical components such as optical filter, beam-splitter and detector, and the retrieval strategies such as *a priori* profiles, spectroscopic MWs, and spectral iterative method are different, which could result in discrepancy in NIR and MIR N<sub>2</sub>O retrievals. To properly compare the NIR and MIR N<sub>2</sub>O retrievals, we use the methodology of Rodgers and Connor (2003) to conform their differences in *a priori* information and AVKs. Both NIR and MIR N<sub>2</sub>O total columns associated with the *a priori* information and the true atmospheric state can be expressed as,

$$TC_r = TC_a + A(PC_t - PC_a) + \varepsilon, \quad (3)$$

where  $A$  represents the AVKs associated with the NIR or MIR retrievals;  $PC_a$  and  $PC_t$  represent *a priori* and true N<sub>2</sub>O partial columns, respectively;  $TC_r$  and  $TC_a$  represent retrieved and *a priori* N<sub>2</sub>O total columns, respectively. The  $\varepsilon$  represents characterization error. Therefore, the discrepancy in N<sub>2</sub>O total columns between MIR ( $TC_{N_2O,MIR}$ ) and NIR ( $TC_{N_2O,NIR}$ ) observations can be expressed as,

$$TC_{N_2O,NIR} - TC_{N_2O,MIR} = (X_{NIR}^{apriori} - X_{MIR}^{apriori}) + (A_{NIR} - A_{MIR})X_{true} + (A_{NIR}X_{NIR}^{apriori} - A_{MIR}X_{MIR}^{apriori}), \quad (4)$$

where  $X_{NIR}^{apriori}$  and  $X_{MIR}^{apriori}$  are *a priori* profiles of N<sub>2</sub>O for NIR and MIR observations, respectively;  $X_{true}$  is the  $X_{N_2O}$  in the true atmospheric state;  $A_{NIR}$  and  $A_{MIR}$  are the AVKs of the retrievals in the NIR and MIR ranges, respectively. As a result, in addition to their different sensitivities of forward model to the real atmospheric state in different spectroscopic MWs, the discrepancy in N<sub>2</sub>O total columns between MIR and NIR observations are dominated by the following two factors: (a) the difference in AVKs and (b) the difference in *a priori* profiles. To reconcile these differences, we use the methodology of Rodgers and Connor (2003) to project the NIR dataset into MIR retrieval scenarios. We first interpolate the NIR daily mean profiles of N<sub>2</sub>O into the MIR vertical height to obtain a common height grid. We then use the *a priori* profiles to correct the interpolated NIR profiles by,

$$X'_{NIR} = X_{NIR} + (A_{NIR} - I)(X_{NIR}^{apriori} - X_{MIR}^{apriori}), \quad (5)$$

where  $X'_{NIR}$  is a N<sub>2</sub>O dataset that has been corrected by the *a priori* profile,  $X_{NIR}$  is the original NIR N<sub>2</sub>O dataset, matrix  $I$  is the unity diagonal matrix. Finally, we further correct  $X'_{NIR}$  using the smoothing function by,

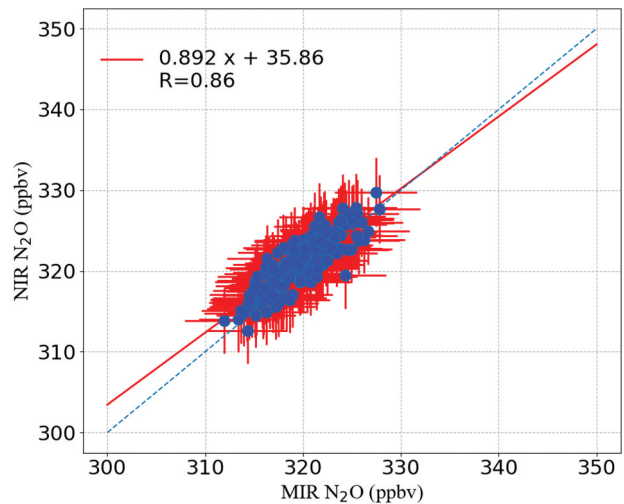
$$X''_{NIR} = X_{MIR}^{apriori} + A_{NIR}(X'_{NIR} - X_{MIR}^{apriori}), \quad (6)$$

where  $X''_{NIR}$  represents a N<sub>2</sub>O dataset that has been corrected by *a priori* profile and smoothing function. This characteristic allows direct comparisons of column observations with in situ data in different seasons.

Following above procedures, all NIR N<sub>2</sub>O data have been projected into MIR retrieval states and compared with the MIR data. Comparisons between daily mean NIR measurements and MIR measurements are shown in Figure 4. We can see in Figure 4 that NIR and MIR  $X_{N_2O}$  datasets are in good agreement, with an average difference of  $(1.33 \pm 4.05)$  ppbv ( $TC_{N_2O,NIR} - TC_{N_2O,MIR}$ ) and a correlation coefficient ( $R$ ) of 0.86. Seasonal cycles of NIR and MIR  $X_{N_2O}$  data in Figure 5 shows that the largest difference of the two datasets occurs in August with a value of  $(-3.11 \pm 5.36)$  ppbv and the lowest difference occurs in December with a value of  $(-0.36 \pm 6.71)$  ppbv ( $TC_{N_2O,NIR} - TC_{N_2O,MIR}$ ). The monthly mean differences between NIR and MIR  $X_{N_2O}$  data from 2015 to 2020 are within  $(1.40 \pm 3.12)$  ppbv. Considering the  $X_{N_2O}$  amplitude is seasonal dependent, we also investigated the variation of the fractional difference between NIR and MIR datasets, which are deduced as the ratios of the absolute discrepancies ( $TC_{N_2O,NIR} - TC_{N_2O,MIR}$ ) to the monthly mean values of the MIR dataset. The results show that the fractional differences of the two datasets are seasonal independent. Therefore, in present work, we first plus a mean bias of 1.4 ppbv to all NIR data and then combine the NIR and MIR datasets to evaluate the seasonal and inter annual variabilities of N<sub>2</sub>O.

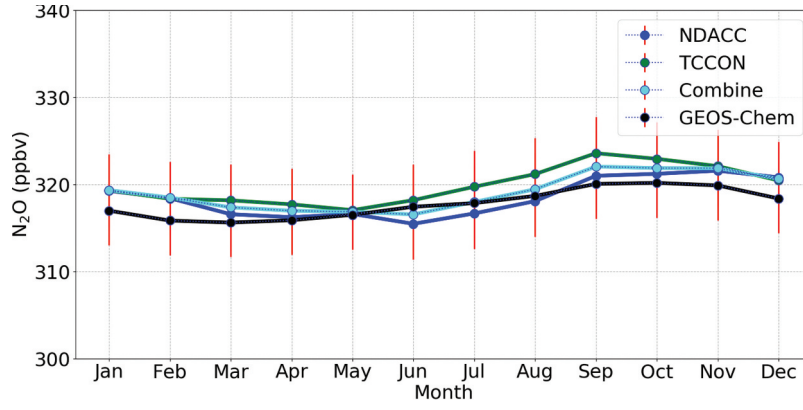
### 3.2. Seasonality and inter-annual trend

In Figure 6, we have used a second Fourier series plus a linear curve to fit the joint NIR+MIR daily mean time series of  $X_{N_2O}$ . Such fitting is based on Gardiner's bootstrap resampling methodology which has been



**Figure 4.** Comparisons between daily mean NIR  $X_{N_2O}$  measurements against MIR  $X_{N_2O}$  observations. The red line represents a linear fit to concurrent scatter points.





**Figure 5.** Monthly mean of NIR, MIR, joint NIR+MIR, and GEOS-Chem model  $X_{N_2O}$  data. The vertical error bars represent the 1- $\sigma$  standard variations.

used in many studies to determine the seasonality and inter-annual trends of many constituents in the atmosphere (Sun et al. 2021c; Yin et al. 2019, 2020, 2021a, 2021b, 2022a). In present work, the measured daily mean time series of  $X_{N_2O}$  related to the regression model is formulated as Equation (7),

$$V(t, \mathbf{b}) = b_0 + b_1 t + b_2 \cos\left(\frac{2\pi t}{365}\right) + b_3 \sin\left(\frac{2\pi t}{365}\right) \quad (7)$$

$$F(t, \mathbf{b}) = V(t, \mathbf{b}) + \varepsilon(t) \quad (8)$$

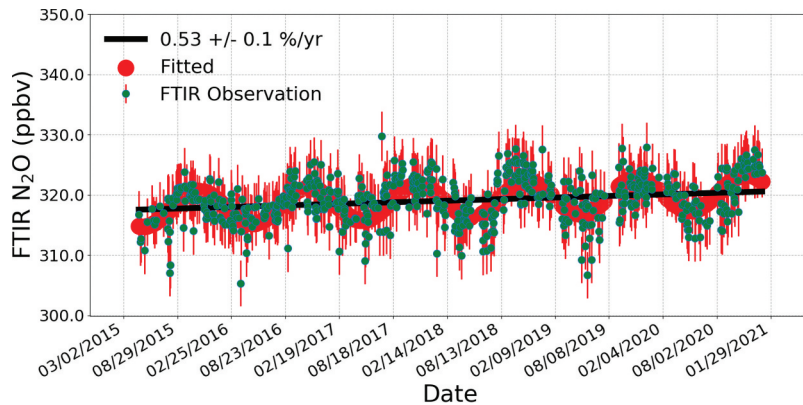
$$d\% = \frac{F(t, \mathbf{b}) - V(t, \mathbf{b})}{V(t, \mathbf{b})} \times 100\%, \quad (9)$$

where  $V(t, \mathbf{b})$  and  $F(t, \mathbf{b})$  are the fitted and measured daily mean time series of  $X_{N_2O}$ , respectively.  $b_0$ ,  $b_1$ ,  $b_2$ , and  $b_3$  are the bootstrap resampling regression coefficients fitted with the model  $V(t, \mathbf{b})$ .  $b_0$  represents the intercept,  $b_1$  is the annual change rate, and  $b_1/b_0$  is the inter-annual trend discussed below.  $b_2$  and  $b_3$  characterize the seasonality,  $t$  represents the fractional of the year elapsed since the start year of the time series (i.e. 2015 in present work), and  $\varepsilon(t)$  is the difference between measured and fitted time series. Equation (9) that calculates the fractional differences of  $X_{N_2O}$  measurements relative to the

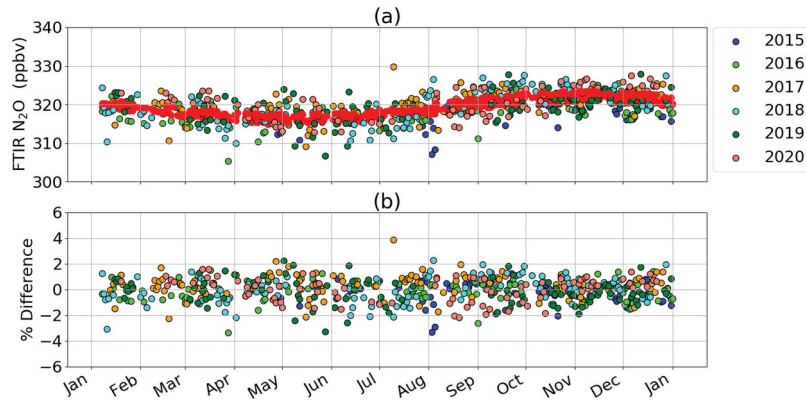
monthly mean values of  $F(t, \mathbf{b})$  is used to analyze seasonal enhancements.

The bootstrap resampling model can generally reproduce the measured seasonality and inter-annual trend of  $X_{N_2O}$  from 2015 to 2020 with an  $R$  of 0.81. Figure 5 showed clear seasonal characteristics of  $X_{N_2O}$  over Hefei: (1) high values of  $X_{N_2O}$  appear in the late summer to early winter and low values of  $X_{N_2O}$  appear in late winter to summer; (2) the variations in  $X_{N_2O}$  are seasonal-independent and relatively stable. (3) the seasonality of  $X_{N_2O}$  varies over a unimodal mode, i.e. the seasonal trough appears around June–July and the seasonal peak appears around September to November.

The joint NIR+MIR  $X_{N_2O}$  time series minimize in June and maximize in September, with values of  $(316.55 \pm 12.22)$  ppbv and  $(322.05 \pm 12.93)$  ppbv, respectively. The joint NIR+MIR  $X_{N_2O}$  measurements in September were on average  $(1.71 \pm 5.22)$  % higher than those in June. The annual mean values of joint NIR+MIR  $X_{N_2O}$  over Hefei are  $(317.90 \pm 5.04)$  ppbv,  $(318.41 \pm 4.90)$  ppbv,  $(319.91 \pm 4.11)$  ppbv,  $(319.74 \pm 5.82)$  ppbv,  $(319.80 \pm 6.21)$  ppbv, and  $(320.83 \pm 3.78)$  ppbv in 2015, 2016, 2017, 2018, 2019, and 2020, respectively. The observed  $X_{N_2O}$  measurements from 2015 to 2020 showed a positive inter-annual trend of  $(0.53 \pm 0.10)$  % per year (Figure 6). Depending on



**Figure 6.** Daily mean time series of joint NIR + MIR  $X_{N_2O}$  measurements between 2015 and 2020 as well as the fitted seasonality and inter-annual trend.

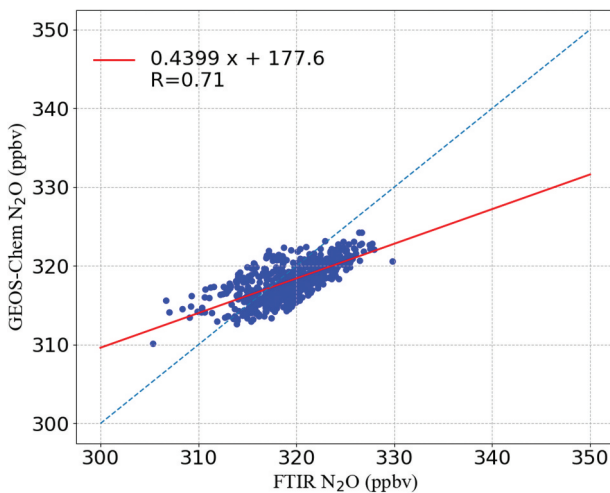


**Figure 7.** (a) Seasonal variabilities of  $X_{N_2O}$  over Hefei based on joint NIR+MIR  $X_{N_2O}$  measurements from 2015 to 2020. The red dots represent monthly mean values of  $X_{N_2O}$ . (b) Fractional differences of  $X_{N_2O}$  measurements relative to the monthly mean values represented by the red dots in (a).

season and observation time, the seasonal  $X_{N_2O}$  enhancements spanned a narrow range of  $-3.89\%$  to  $4.05\%$  (Figure 7). The small values of inter-annual trend and seasonal enhancement are mainly attributed to the chemical stable characteristics of  $N_2O$  in the atmosphere.

### 3.3. Comparisons with GEOS-Chem model

GEOS-Chem simulations within  $\pm 1$  hour of the FTIR measurement times are compared with the joint NIR + MIR observations. Similar to section 3.1, we also use the methodology of Rodgers and Connor (2003) to project the GEOS-Chem model dataset into MIR retrieval scenarios before using them in comparison. The GEOS-Chem simulations sample at the nearest ground pixel to the Hefei site are first vertically interpolated into the MIR vertical height. The interpolated data are then smoothed with the AKs and a priori profiles of MIR retrievals via Equation (10).



**Figure 8.** Correlation plots of the joint NIR+MIR data against GEOS-Chem model data. The red line represents a linear fit to the scatter points. The dotted grey line is the 1:1 line.

$$X_s = X_a + A(X_c - X_a), \quad (10)$$

where  $X_c$  and  $X_s$  are the interpolated and smoothed GEOS-Chem profile, respectively;  $A$  and  $X_a$  are AK matrix and a priori profile of MIR retrievals, respectively. The GEOS-Chem  $X_{N_2O}$  is obtained by an integration of the smoothed GEOS-Chem profile from ground to the 120 km.

GEOS-Chem model  $X_{N_2O}$  data were generally consistent with FTIR  $X_{N_2O}$  data over Hefei, with a correlation of 0.71 (Figure 8). However, GEOS-Chem model data tend to underestimate  $N_2O$  concentration and its seasonality in Hefei, especially in autumn and winter (Figure 5). The maximum and minimum values of GEOS-Chem model  $X_{N_2O}$  data are  $(320.17 \pm 1.31)$  ppbv and  $(315.61 \pm 1.52)$  ppbv in November and March, respectively. Monthly mean differences between FTIR measurements and GEOS-Chem data are within  $(1.34 \pm 3.12)$  ppbv, where the maximum and minimum differences are  $(2.65 \pm 2.14)$  ppbv and  $(0.88 \pm 2.05)$  ppbv in February and June, respectively. The discrepancies between GEOS-Chem simulations and FTIR measurements are mainly attributed to uncertainties of model inputs such as emission inventories and meteorological fields.

## 4. Conclusions

Both NIR and MIR solar spectra saved by ground-based high resolution FTIR spectrometer were used to retrieve  $X_{N_2O}$  over Hefei, eastern China. The  $X_{N_2O}$  retrievals derived from NIR and MIR solar spectra were characterized and compared. Generally, NIR and MIR measurements agree well with an  $R$  of 0.86 and an average difference of  $(1.33 \pm 4.05)$  ppbv (NIR - MIR). By reconciling this difference, the joint NIR+MIR observations disclose that the monthly mean time series of  $X_{N_2O}$  minimize in June and maximize in September, with values of  $(316.55 \pm 12.22)$  ppbv and  $(322.05 \pm 12.93)$  ppbv, respectively. The  $X_{N_2O}$  measurements in September were on average  $(1.71 \pm 5.22)\%$  higher than

those in June. The observed  $X_{N_2O}$  measurements between 2015 and 2020 showed a positive inter-annual trend of  $(0.53 \pm 0.10)$  %/year. We also compared GEOS-Chem simulations with the joint NIR+MIR measurements. GEOS-Chem model  $X_{N_2O}$  data were generally consistent with FTIR measurement  $X_{N_2O}$  data over Hefei, with a correlation coefficient and an average difference of 0.71 and  $(1.25 \pm 4.31)$  ppbv, respectively. This study can enhance current knowledge of ground-based high-resolution FTIR remote sensing of  $N_2O$  in the atmosphere and contribute to generating a new reliable  $N_2O$  dataset for climate change research.

## Disclosure statement

No potential conflict of interest was reported by the author(s).

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## Data availability statement

Data underlying the results presented in this paper are not publicly available at this time but may be obtained from the authors upon reasonable request.

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